

# A noise Power Indicator ... by Brian Coleman, G4NNS

A noise power indicator can be a useful tool for judging the performance of microwave receivers using the Y factor method (cold sky vs warm object). This relies on the fact that all bodies above absolute zero emit "black body" radiation which can be detected at the higher microwave frequencies. In practice this means that you can use the "cold" sky vs a known "warm" reference. I find aiming the antenna at the house nearby is best and you can assume its temperature is about 293K. Measurement as opposed to indication can be accomplished by using a precision attenuator in front of the indicator and using the same level at the detector to avoid any non-linearity in the system. I also use my system to ensure that my EME antenna is tracking correctly using moon noise of which I see about 2.5dB at 10GHz. Other uses of this tool can include antenna testing. It provides a more precise and easier to read system than an S-meter and can, in conjunction with a precision attenuator provide a means of making fairly precise measurements of the order of +/- 0.2dB. When using a warm source such as your house remember that it should fill the aperture of your antenna. This is why small antennas (more than about 0.5 degrees beam width) do not see the full potential sun or moon noise.

## Circuit description (see Figure 1)

The system consists of a high gain amplifier (circa 70dBs) and a sensitive detector at the IF frequency (in this case 144MHz). The bandwidth needs to be as wide as possible but not exceeding that of the transverter front end. This is important as noise generated outside the desired pass band of the transverter is not relevant to the system performance and could swamp the power indication and lead to misleadingly poor indication of receiver performance. The 3dB bandwidth of this indicator is about 1MHz. Relative measurements are made by using a switched attenuator in front of the noise power indicator. Such attenuators can be found at microwave flea markets.

The main challenges to the circuit design and construction are:- 1) minimising the gain required to drive the detector. I.e. the detector must be as sensitive as possible and 2) keeping the gain stages stable at all frequencies including those outside the intended pass band. The first requirement is met by using forward bias on the schottky detector diode to move it into its linear and most sensitive region and by using a simple op-amp to amplify the DC difference between the RF detector and another matched reference diode, similarly forward biased but decoupled to RF, in the same SM package. The second requirement is achieved by using MMIC gain stages with careful construction, grounding and shielding. With so much gain on one frequency, great care has to be taken in the layout and construction to avoid instability. The stage before the detector and second band pass filter uses a JFET as this is more tolerant of impedance mis-match. No originality is claimed for any part of the circuit but some time and effort has been spent in trying to ensure the repeatability of the layout and simplicity of the PCB. If good practice for RF construction is followed, it should be possible to build the circuit and get it to be stable quite easily. The input must however look substantially resistive and be correctly matched at all frequencies. This can be achieved by the use of some fixed attenuation at the input.

## Construction (see Figure 2)

A simple way to make the PCB is to cover the double sided copper laminate with sticky tape such as 3M Magic tape then stick a print of the track layout onto the tape and use a scalpel or sharp modelling knife to cut out and peel off the areas to be exposed to the etchant.

The design uses as many straight lines as possible to simplify this. I found it helpful to use a straight edge under a 3D magnifier. After etching and cleaning the PCB, the screens should be cut to size. The long, side screens are best made using 1/16" laminate as used for the main PCB. The inter stage screens are best made from 1/32" copper laminate or from tin plate. I did not find end screens to be necessary but if you fit them they should be made from 1/16" laminate. Drill 2.5mm holes to recess the MAR8s, which should be the "drop in" type, to minimise lead lengths. Also drill 0.8mm holes for the Via pins. These pins are essential for stability and it is possible that more than the four shown may be needed. (see setting up).

As with any PCB construction it is best to start with the smallest components first as the larger ones will otherwise get in the way (see figures 3 & 4). Whilst the screens should be fitted last you may find it easier to fit L1 and L2 after fitting the long side screen closest to these inductors as this will make it easier to solder the seam. L1 and L2 are separated by about 3mm and are 2-3mm above the ground plane. Note also that the long side screens extend below the main PCB and should match the height of the inter-stage screens on the component side. They should be soldered both to the top and bottom surfaces of the main PCB. Make sure the side screens remain at 90 degrees to the PCB by "tack soldering" both sides and repositioning as necessary before soldering the entire length of the seams.

When assembling the DC amplifier it is again best to fit the smallest components including the wire links and pins first. I find it best to make the track cuts after the components are in place and to use this process as a check

for correct placement of the components. Start fitting the active components with the +5V regulator, and test this before fitting the voltage converter. Test that the + and - 10V rails are working before fitting the op-amp.

### Setting up

Terminate the input with a good 50 Ohm load. This is essential for stability. Adjust VR1 for zero reading on the meter. If this is not possible the amplifier may be oscillating. This should not be happening! If it is try touching different gain stages to see if the meter reading changes. You may be able to identify which stage is oscillating by this method. You could try lifting one end of each one Ohm resistor in turn to identify the oscillating stage. Whatever you try, make sure it is reversible and restore it to the original condition before trying something else. Here are some things to try if necessary: 1) Change the capacitor and inductor tuning to see if the oscillation stops. 2) Add additional via pins close to the MAR8s. Those indicated in the design have proved sufficient but in some cases more may be required to achieve stability. 3) Try adjusting the values for R1, R6 or R7 for the stage that is oscillating.

Once the amplifier is stable and the meter set to zero, apply a small signal within the pass-band e.g. 144.50MHz from a 50 Ohm source through a switched attenuator. Tune VC1, VC2, L3 and L4 for maximum reading increasing the attenuation as you go to keep the meter on scale. The prototypes, when set up as described, had a Full Scale reading on the most sensitive range, corresponding to about -80 to -85dBm.

Next, connect to the receiver or transverter under test. The Indicator should show a non zero reading when the transverter is switched on but make sure there are no strong signals in the pass-band, such as local beacons. If the meter is on the end stop try inserting some attenuation between the transverter and the indicator

### Applications

Receiver performance can be measured using ground noise and quiet sky noise. For EME systems where the antenna beam width is close to or less than the angular diameter of the source, Sun and moon noise measurements provide a useful indication of receive performance. The indicator can also be used for calibration tracking systems. The indicator also has potential for plotting polar diagrams and measuring gain of antennas for 2m or bands where a transverter has a 2m IF.

For EME, the indicator can display moon noise during receive periods to provide confirmation that the tracking system is functioning correctly.

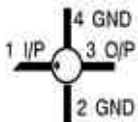
For this purpose, it is necessary to supply a sample of the IF to the indicator through a splitter ( T ) and to switch the indicator out of circuit during transmit periods. It is desirable to switch off some of gain stages during transmit periods, to avoid overload. Some gain drift may be experienced but this is minimised if the smallest number of stages are switched off. Try switching off just the first one or two stages if possible. A spare power track is supplied on the PCB for this purpose. The circuit diagram shows the possible arrangement for RF switching. The number of stages to be switched will depend on the isolation of the relay used.

### Notes

1. If you fit the unit in a box and it becomes unstable, try lining the box with conducting foam or lossy rubber where it is closest to the component side of the Noise amplifier.
2. On the supplied veroboard component layout for the DC amplifier, **C26** is missing. This decoupling capacitor can be conveniently soldered across the MX680 IC, from pin 8 to pin 5.
3. Further information, if required can be obtained from Brian, G4NNS, QTHR or

email to: [brian-coleman@tiscali.co.uk](mailto:brian-coleman@tiscali.co.uk)

4. A component list, **figure 4**, is provided at the end of this article.
5. Except for the MAR8 modamps, semiconductor pin outs are shown on the circuit diagram.



**MAR8 pin out connections**

Figure 1

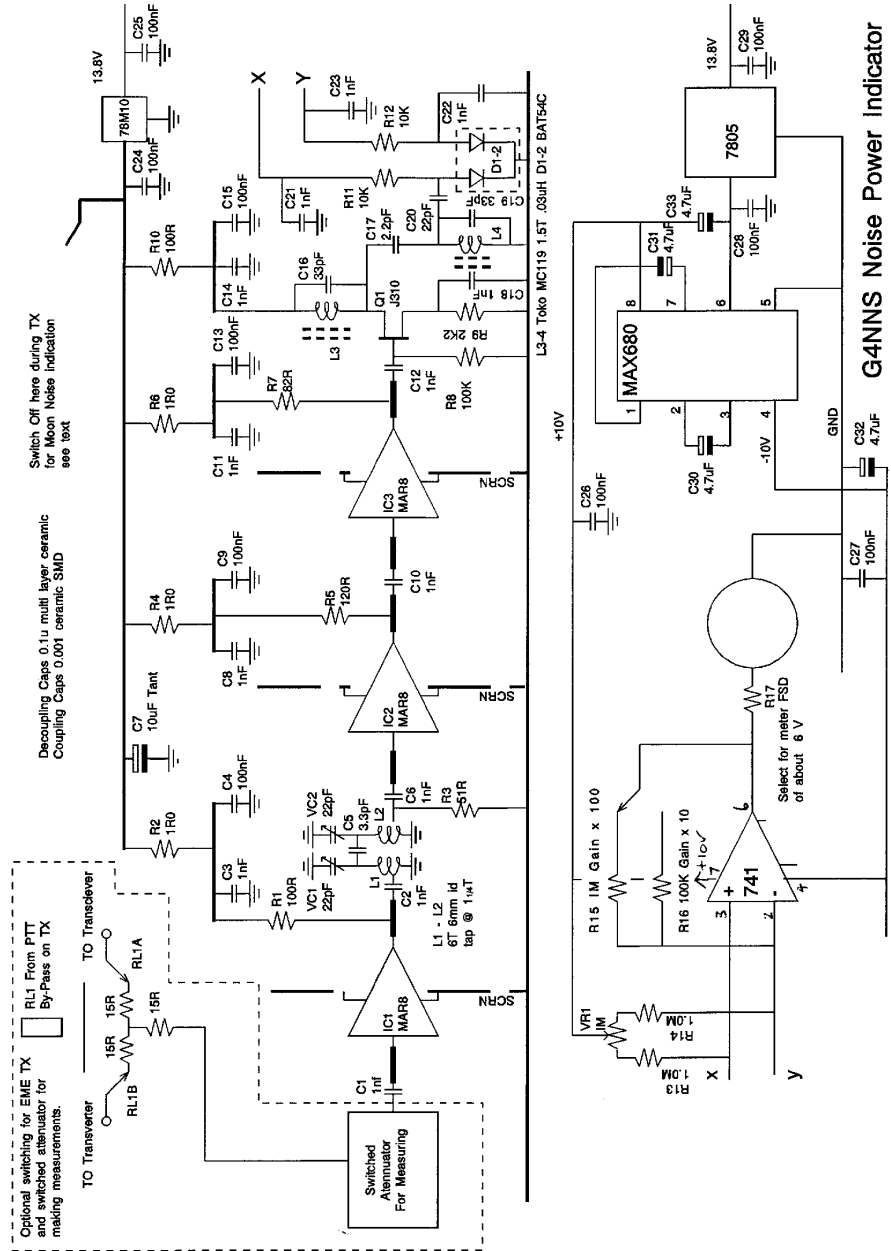
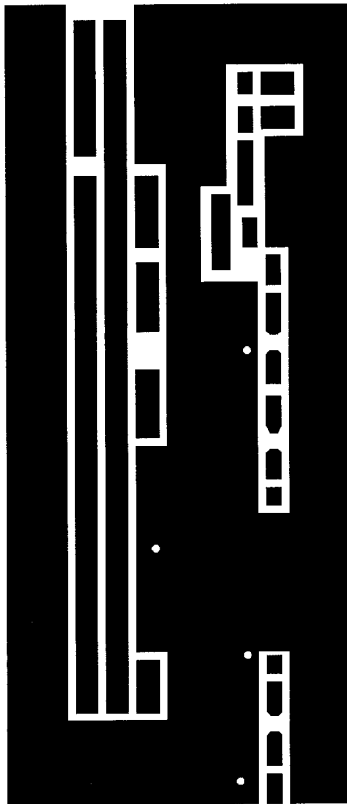


Figure 2



Main PCB 1/16 Laminate 105 x 45mm  
Note position of 4 via holes.

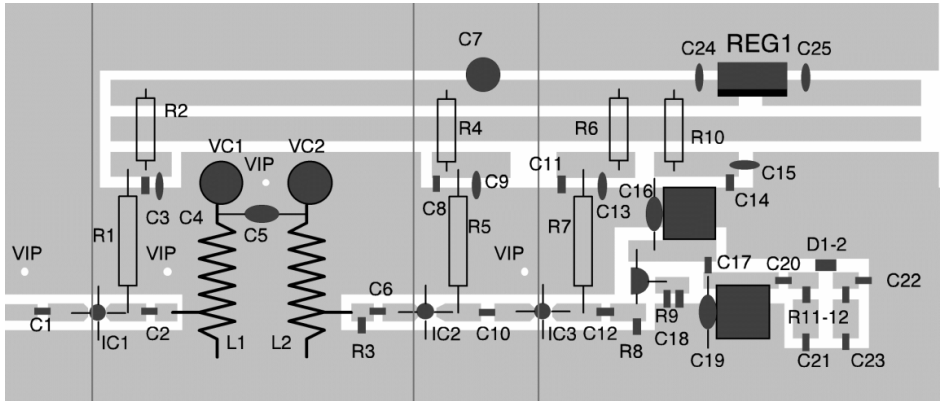


Intermediate Screens  
1/32" Laminate  
45 x 20mm  
QTY 3



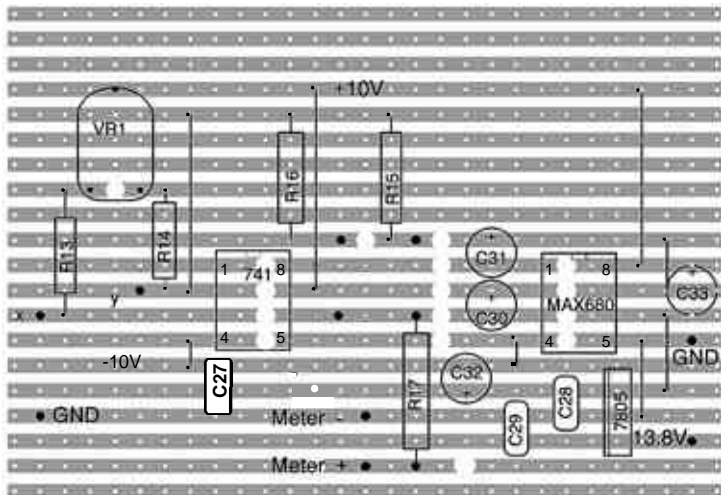
Side screens 1/16" Laminate 105 x 25mm . . . QTY 2

# G4NNS NOISE AMPLIFIER (continued)



Above: Main amplifier pcb tack and component layout

Below: DC amplifier Veroboard layout. Note that C26 is missing from this diagram and should be soldered from the +10V line to ground at a convenient point (eg across pins 5 and 8 of the MAX680)



## G4NNS NOISE AMPLIFIER—COMPONENT LIST

R1	100R Ax		C1	1nF SM
R2	1R Ax		C2	1nF SM
R3	51R SMD		C3	1nF SM
R4	1R Ax		C4	100nF SM
R5	120R Ax		C5	3.3pF Rad
R6	1R Ax		C6	1nF SM
R7	82R Ax		C7	10uf Tant
R8	100K SM		C8	1nF SM
R9	2K2 SM		C9	100nF SM
R10	100R Ax		C10	1nF SM
R11	10K SM		C11	1nF SM
R12	10K SM		C12	1nF SM
R13	1M0 Ax		C13	100nF SM
R14	1M0 Ax		C14	1nF SM
VR1	1M		C15	100nF SM
R15	1M Ax		C16	33pf Rad
R16	100K Ax		C17	2.2pF SM
R17	SOT		C18	1nF SM
			C19	33pF Rad
VR1	1M0		C20	22pF SM
			C21	1nF SM
VC1	22pF		C22	1nF SM
VC2	22pF		C23	1nF SM
			C24	100nF SM
L1	6T 6mm id		C25	100nF SM
L2	6T 6mm id		C26	100nF Rad
L3	MC119 1.5T	1st 10 SAE to G4NNS	C27	100nF Rad
L4	MC119 1.5T	or try BONEX	C28	100nF Rad
			C29	100nF Rad
			C30	4.7 uF
			C31	4.7 uF
			C32	4.7 uF
			C33	4.7 uF
Q1	J310			
IC4	741			
IC5	MAX 680	Farnell 246_451		
REG1	7810			
REG2	7805			
IC1	MAR8			
IC2	MAR8			
IC3	MAR8			
D1-2	BAT 54C	Farnell 302_284		